



Palaeokarstic surfaces in the Upper Cretaceous limestones of central Jamaica

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Abstract

Two prominent pitted surfaces in rhythmic, late Maastrichtian, rudist-bearing limestones in central Jamaica are described. The lower surface is pock-marked with small pits that form a nodular, mamillated surface. The pits are roughly circular to gently elliptical in plan view. They vary in diameter from 5–170 mm and in depth from 5–90 mm, and are orientated perpendicular to bedding. This surface is buried beneath a lignite-rich siltstone bed. The upper surface contains microkarren phenomena represented by small cylindrical potholes up to 40 mm in diameter and 30 mm deep. These surfaces are interpreted as palaeokarstic because the pits are: (1) perpendicular to bedding and not the present land surface; (2) infilled with very poorly sorted siltstone, and (3) overlain by lignitic shales indicating terrestrial influx. The presence of two palaeokarstic surfaces in late Maastrichtian limestones suggests short-period cycles in relative sea level, consistent with glacial-eustatic driving mechanisms.

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Keywords: Maastrichtian; limestones; palaeokarst; rudists; Jamaica.

1. Introduction

Palaeokarst can be defined as ‘fossil’ karst, a remnant from a previous period of karstification, and characterised by the presence of ancient, buried surfaces and deposits. It is distinguished from relict karst, which originated within the present phase of karstification, and although the geomorphic processes which formed the terrain are no longer operative, the karst is still being modified by contemporary processes. Palaeokarst is, therefore, buried and inert, and although it can include karst of interstratal and hypogene origin (Ford & Williams, 1989), most documented palaeokarsts are surficial, epikarstic phenomena and cavernous features (Bosak et al., 1990). Palaeokarst occurs across the same range of spatial scales as does active karst, from small-scale solution sculpturing to surface features with tens to hundreds of metres of relief (Ford & Williams, 1989). At the small extreme, one of the most common types of palaeokarst is solution pitted exposure surfaces capping

shallow-water limestones forming cycles or rhythms, where the carbonates were briefly exposed to vadose diagenesis and karren pitting before re-submergence and burial by new carbonate deposition (Ford & Williams, 1989).

The Late Cretaceous (Maastrichtian) to early Palaeocene succession in the Central Inlier (central Jamaica; Fig. 1) is represented by the Kellits Synthem (Mitchell, in press). This shows a transgressive-regressive cycle of probable Caribbean-wide tectonic significance (Mitchell & Blissett, 2001; Mitchell, in press). The transgressive succession begins with braided-stream type conglomerates, passes up through tidal flat sandstones-mudstones and into thick rudist-bearing limestones (Mitchell, 1999; Mitchell & Blissett, 2001; Mitchell, in press). The regressive succession consists of a volcanoclastic, braid-delta complex (Mitchell, 2000; Mitchell & Blissett, 2001). Although there have been extensive studies of the palaeontology of the limestone of the Kellits Synthem (e.g., Whitfield, 1897; Trechmann, 1924; Chubb, 1971; Coates, 1977; Jiang & Robinson, 1977; Sohl & Kollman, 1985; Donovan, 1993; Hazel & Kamiya, 1993; Krijnen et al., 1993; Sandy et al., 1997; Mitchell, 2002a,b;

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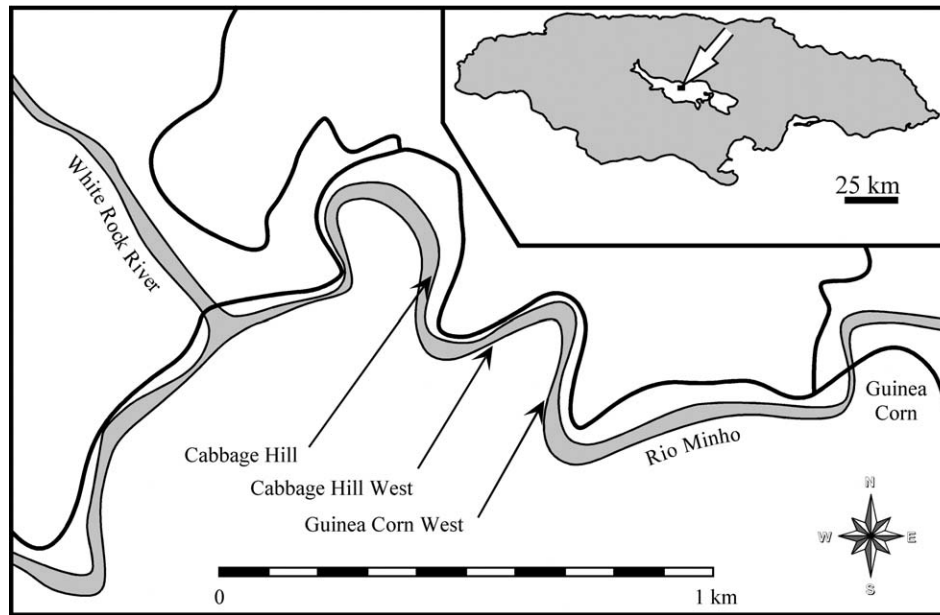


Fig. 1. Distribution of sections studied in the Central Inlier of central Jamaica. Rocks shown as black lines, rivers as grey lines, localities arrowed. The inset map shows the localities of the Central Inlier (unshaded) and the position of the main map (black box).

62 Mitchell & Gunter, 2002), there have been few studies of
 63 the sedimentology. In this paper we document two
 64 recently discovered palaeokarstic surfaces within the
 65 Guinea Corn Formation.

66 2. Rhythms in the upper Guinea Corn Formation

67 The Guinea Corn Formation was deposited on a
 68 shallow shelf/platform adjacent to an active andesitic
 69 volcanic complex (Mitchell, 2002a,b). These shallow
 70 water limestones can be traced some 100 km across
 71 central and western Jamaica. The succession in central
 72 Jamaica consists of some 200 m of shallow-water, rudist-
 73 bearing limestones and intervening mudstones and
 74 sandstones that are arranged into rhythms (Mitchell,
 75 1999, 2002b). The rudist bearing limestones have
 76 recently been dated as Late Maastrichtian (Steuber
 77 et al., 2002). The rhythms in the upper 50 m, or so, of the
 78 Guinea Corn Formation consist of a lower siltstone or
 79 calcareous siltstone unit, and an upper limestone unit
 80 (Mitchell, 1999, 2002b). The siltstone unit is usually
 81 relatively thin, yields few fossils, and may contain a large
 82 proportion of lignite fragments. Generally, it has a
 83 rather abrupt contact with the underlying limestone at
 84 the top of the previous rhythm. The limestone portion
 85 of each rhythm can usually be divided into two: a
 86 lower part that contains relatively common rudists
 87 [*Titanosarcolites giganteus* (Whitfield), *Biradiolites*
 88 *jamaicensis* Trechmann, and *Thyrastylon* sp.] together
 89 with abundant larger benthic foraminifera [*Chubbina*
 90 *jamaicensis* Robinson and *Kathina jamaicensis* (Jarvis
 91 and Cushman)]; and an upper part in which rudists are

rare, while the same species of benthic foraminifera
 remain abundant. In this paper we consider the succes-
 sion of rhythms in the upper part of the Guinea Corn
 Formation that were labelled F5 to G5 by Mitchell
 (1999).

2.1. Rhythm F5

Rhythms in the F Beds of the Guinea Corn Forma-
 tion are characterised by a relatively thin siltstone/
 sandstone lower division (typically 20–100 cm thick) and
 a thick upper limestone division (typically 80–300 cm
 thick) (Mitchell, 1999). Rhythm F5 is the uppermost
 rhythm of the F Beds. Its basal contact with rhythm F4
 is relatively abrupt (surface 1 in Fig. 2) and shows no
 palaeokarstic features. The limestone contains a low
 diversity of rudist bivalves together with the benthic
 foraminifera *K. jamaicensis* and *C. jamaicensis* (Fig. 2).

2.2. Rhythms G1–G4

These four rhythms can easily be traced between
 exposures in the Guinea Corn Formation (Fig. 2). The
 correlation of rhythms is aided by the distinctive nature
 of rhythm G3, a grey limestone (the other rhythms
 comprise white limestones) containing abundant white
 larger benthic foraminifera and common steinkerns of
 gastropods.

The base of rhythm G1 is particularly noteworthy as
 it consists of a prominent irregular surface overlain
 locally by a lignite-rich mudstone unit (=surface 2 in
 Fig. 2). No rootlets are present beneath the lignitic shale.
 At Cabbage Hill, the surface occurs on the upper

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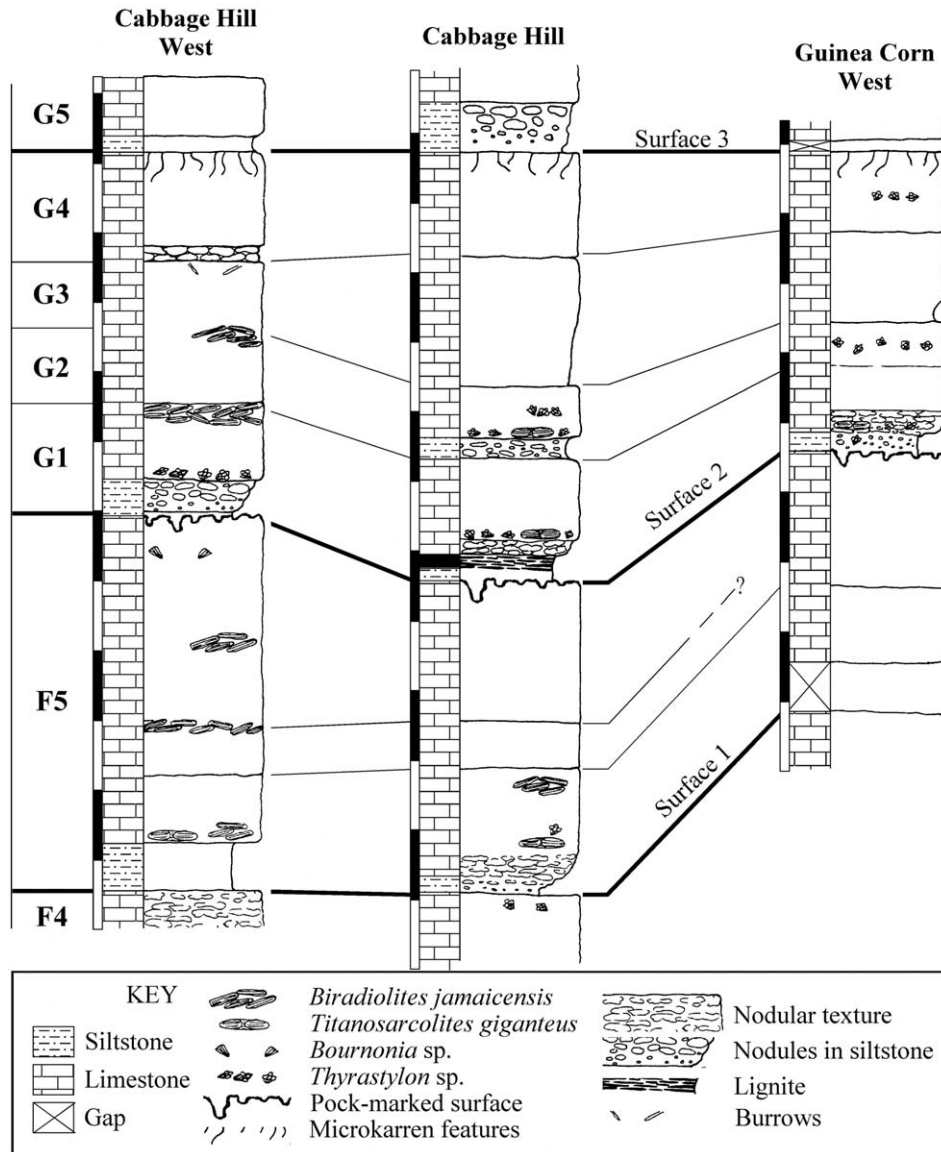


Fig. 2. Graphic logs showing three sections through rhythms F4 to G5 of the Guinea Corn Formation (rhythm numbers after Mitchell, 1999), at the three localities shown in Fig. 1, using the base of rhythm G5 as datum. Graphic logs show weathering profiles. Rudist abundance recorded as seen in the field. Surfaces 1–3 (2–3 are palaeokarstic) are described in detail in the text. Scale bars in metre intervals.

121 bedding plane of the steeply dipping (c. 70 degrees),
 122 fairly massive uppermost-limestone bed of rhythm F5,
 123 and displays a small-scale joint pattern normal to
 124 bedding (Fig. 3A). The joint pattern has been opened up
 125 slightly by solutional weathering, forming a rubbly and
 126 nodular texture that extends beneath the surface to
 127 a depth of 41 cm. Distinct honeycomb weathering
 128 patterns, similar to spitzkarren, with a predominant
 129 orientation normal to bedding, can also be detected to a
 130 similar depth; these are partly infilled with the sandy
 131 mudstones of the overlying bed. The surface is pock-
 132 marked with small pits that form a nodular, mamillated
 133 surface (Fig. 3B). Most of the pits are roughly circular to
 134 gently elliptical in plan view, although a few are more
 135 irregular. The pits vary in diameter from 5–170 mm and

136 in depth from 5–90 mm, and are orientated perpendicu-
 137 lar to the overlying bedding surface. The features can be
 138 divided into two predominant types based on mor-
 139 phometry (Table 1). The most common karren features
 140 are pits and pans, which are roughly circular to oval in
 141 plan view, with bowl-shaped floors, 45–170 mm in diam-
 142 eter and 25–90 mm deep. The larger pits commonly
 143 have steep boundary walls that slope convexly inwards
 144 towards the rounded floor of the pit, making them
 145 convexo-concave in profile (Fig. 3C). Some of the larger
 146 pits are partly infilled with sediment similar to the
 147 overlying bed, while others are devoid of any infill. A
 148 few of the pits have flattened floors and the bounding
 149 slopes have been steepened by undercutting, forming a
 150 distinct break of slope between pit floor and side slope,

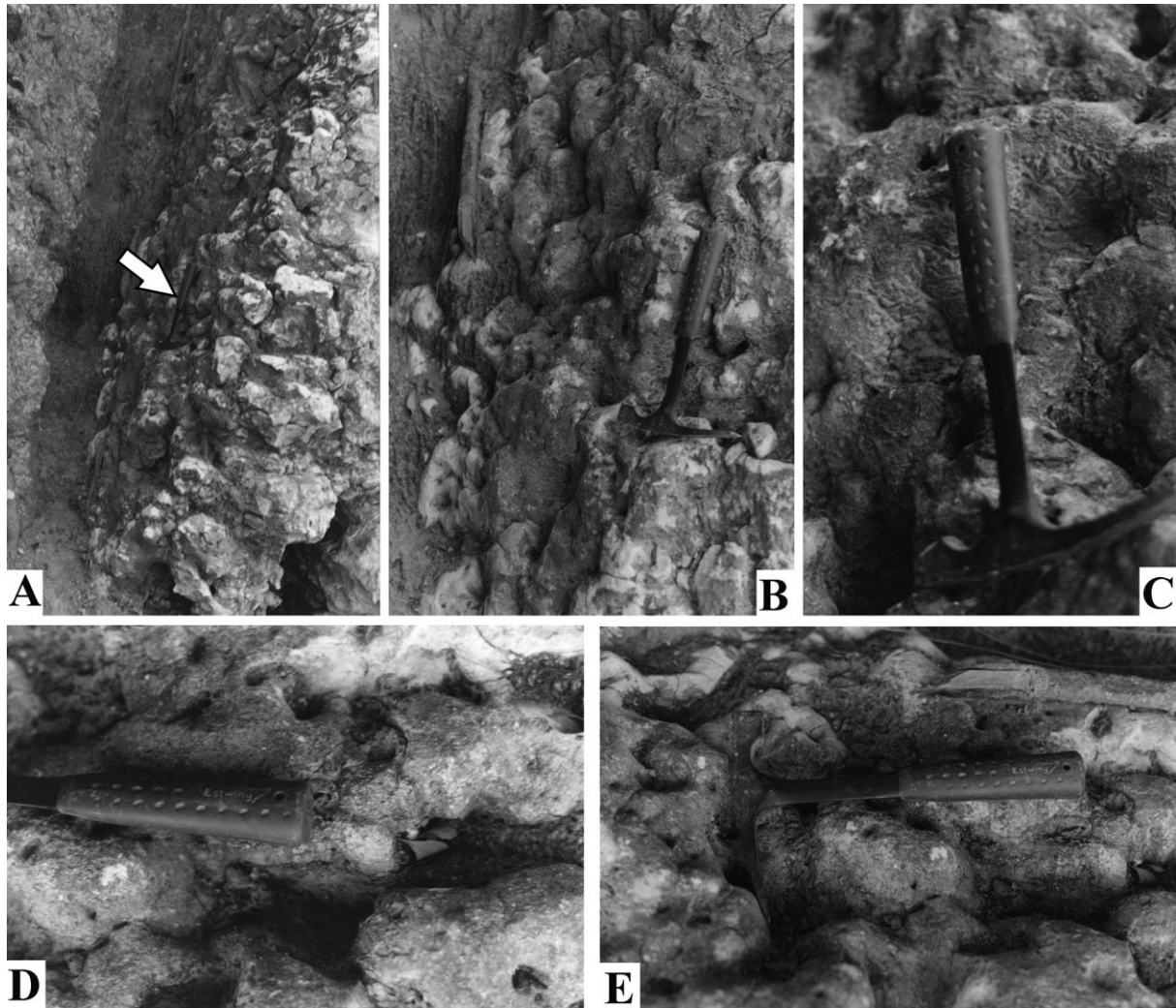


Fig. 3. Photographs of palaeokarstic surface 2 in the Guinea Corn Formation, Jamaica. A, palaeokarstic surface at Cabbage Hill, surface 2, cycle G1–4, showing steeply dipping nature of the cycle, with overlying sedimentary rocks with abundant lignite. The surface is clearly nodular and mamillated in appearance, with the development of an open joint normal to the palaeokarstic surface. A larger pothole appears at the base of the exposure. Hammer (arrowed) is 28 cm long. B, detail of the palaeokarstic surface at Cabbage Hill, surface 2, clearly showing pitted and uneven surface with larger pits and smaller pock-marks. C, larger bowl-shaped pits on the palaeokarstic surface, showing steep, convex, bounding slopes and rounded pit floors. Some of the pit floors are partly infilled with sedimentary rocks. D, solution pans with steep boundary slopes having been over-deepened and undercut to form small basal notches. The floors of some of the pans are partly infilled with sedimentary rocks. E, larger solution pits with intervening rain splash pits; the latter are superimposed on the positive relief elements of the surface. The smaller rainpits do not develop on the bounding slopes or floors of the larger solution pits and pans.

151 leading to a more distinct pan shape. Indeed, a few of
 152 the pits have over-steepened, almost undercut boundary
 153 slopes that form a small basal notch at the junction
 154 between the pit floor and the bounding wall (Fig. 3D).
 155 The larger pits are clustered and a few are aligned. Some
 156 of the larger pits are elongated, although there is no
 157 common orientation pattern.

158 The surface at the base of rhythm G1 also contains a
 159 few smaller pits and pock-marks, which are super-
 160 imposed upon, and restricted to, the upstanding relief
 161 elements of the surface, forming small imprints in areas
 162 between the larger pits (Fig. 3E). They do not occur
 163 within the larger pits themselves. These smaller features
 164 are also circular to oval in plan, but predominantly

165 tubular to cylindrical rather than bowl-shaped in bed-
 166 normal orientation. Some of the smaller pits are tapered.
 167 They are less common and more isolated than the larger
 168 pits, and range from 5 to 40 mm in diameter and up to
 169 30 mm deep. Most of the smaller pock-marks are devoid
 170 of a sedimentary infill (i.e., it has been removed by
 171 weathering).

172 The basal G1 surface at Cabbage Hill exposes 27
 173 measurable pits of varying size and their overall mor-
 174 phometric properties are given in Table 1. The larger
 175 pits have a tendency to be slightly more elongated, and
 176 have shallower depth/diameter and width/depth ratios
 177 compared to the smaller features, reflecting the bowl-
 178 shaped cross-sections of the larger pits, compared to the

202 Table 1
 203 Morphometric properties of the palaeokarstic surface at Cabbage Hill, Surface 2, Rhythm G1–4 (measurements in mm)

203	Classification	Length (L)	Width (W)	Depth (D)	Mean diameter MD=(L+W)/2	Depth/Diameter =D/MD	Elongation =L/W	Width/Depth =W/D
204	Small pit	5	5	10	5	2.00	1.00	0.50
205	Small pit	15	10	10	12.5	0.80	1.50	1.00
206	Small pit	20	15	15	17.5	0.86	1.33	1.00
207	Small pit	25	25	20	25	0.80	1.00	1.25
208	Small pit	30	30	25	30	0.83	1.00	1.20
209	Small pit	35	30	30	32.5	0.92	1.17	1.00
210	Small pit	35	25	20	30	0.67	1.40	1.25
211	Small pit	40	25	20	32.5	0.62	1.60	1.25
212	Small pit	45	15	25	30	0.83	3.00	0.60
213	Small pit	45	35	30	40	0.75	1.29	1.17
214	Mean	29.5	21.5	20.5	25.5	0.9	1.4	1.0
215	Std dev.	13.2	9.7	7.2	10.7	0.4	0.6	0.3
216	Large pit	50	35	30	42.5	0.71	1.43	1.17
217	Large pit	50	45	35	47.5	0.74	1.11	1.29
218	Large pit	50	35	25	42.5	0.59	1.43	1.40
219	Large pit	55	35	30	45	0.67	1.57	1.17
220	Large pit	60	55	30	57.5	0.52	1.09	1.83
221	Large pit	65	50	25	57.5	0.43	1.30	2.00
222	Large pit	75	45	35	60	0.58	1.67	1.29
223	Large pit	80	65	40	72.5	0.55	1.23	1.63
224	Large pit	85	80	70	82.5	0.85	1.06	1.14
225	Large pit	100	65	60	82.5	0.73	1.54	1.08
226	Large pit	105	75	50	90	0.56	1.40	1.50
227	Large pit	110	70	45	90	0.50	1.57	1.56
228	Large pit	115	60	45	87.5	0.51	1.92	1.33
229	Large pit	120	85	45	102.5	0.44	1.41	1.89
230	Large pit	150	45	40	97.5	0.41	3.33	1.13
231	Large pit	150	75	60	112.5	0.53	2.00	1.25
232	Large pit	240	100	90	170	0.53	2.40	1.11
233	Mean	97.6	60.0	44.4	78.8	0.6	1.6	1.4
234	Std dev.	49.3	19.2	17.4	32.4	0.1	0.6	0.3

179 tubular and cylindrical form of the smaller pits. Fig. 4
 180 shows that there is a distinct relationship between
 181 length-to-width and depth-to-diameter parameters of
 182 the pits. Relationships between other parameters show
 183 no distinct correlations.

184 Two larger and deeper potholes also occur on the
 185 surface with an orientation normal to bedding. The larg-
 186 est pothole occurs towards the base of the exposure (Fig.
 187 3A) and extends through the body of the limestone bed
 188 forming a small ‘lighthouse’. The pothole is 84 cm deep
 189 and 20 cm wide. It is broadly cylindrical in shape and
 190 circular in plan with no apparent elongation. The interior
 191 surface of the pothole is scalloped and marked by small
 192 flutes, suggesting vertical water movements down the
 193 feature. The joint pattern is better developed and has a
 194 more open appearance in the immediate vicinity of the
 195 pothole compared to the remainder of the bed, suggesting
 196 that the origin of the feature was controlled by enhanced
 197 solution along more common lines of weakness. A second
 198 larger, broadly circular, cylinder-shaped pothole also
 199 occurs towards the base of the exposure. This measures
 200 24.5 cm in diameter and is 23 cm deep, and is partly
 201 infilled with sediment from the overlying bed.

235 The sedimentary rock overlying and infilling the
 236 features of the top surface of F5 consists of very poorly
 237 sorted, sandy siltstone. The sand grade material includes
 238 highly weathered andesitic lithic fragments and crystals.
 239 This poorly sorted, sandy siltstone is overlain by a
 240 30 cm-thick lignite bed. The lignite consists of flattened
 241 fragments of wood; rootlets are absent. Neither marine
 242 nor terrestrial body fossils have been recovered from the
 243 lignite bed, only plant remains. The lignite bed is over-
 244 lain by a nodular limestone that contains the benthic
 245 foraminifers *C. jamaicensis* and *K. jamaicensis*.

246 Surface 2 is also well developed at the two other
 247 locations exposing this part of the Guinea Corn Forma-
 248 tion (Fig. 2). The surface in each case shows similar
 249 characteristics to those seen at Cabbage Hill.

250 2.3. Rhythm G5

251 Rhythm G5 marks a thick limestone in the Guinea
 252 Corn Formation and rests on a prominent, distinctive
 253 surface (surface 3 in Fig. 2). At Cabbage Hill, this
 254 surface is pock-marked and highly irregular, although it
 255 is much less nodular and mamillated in appearance

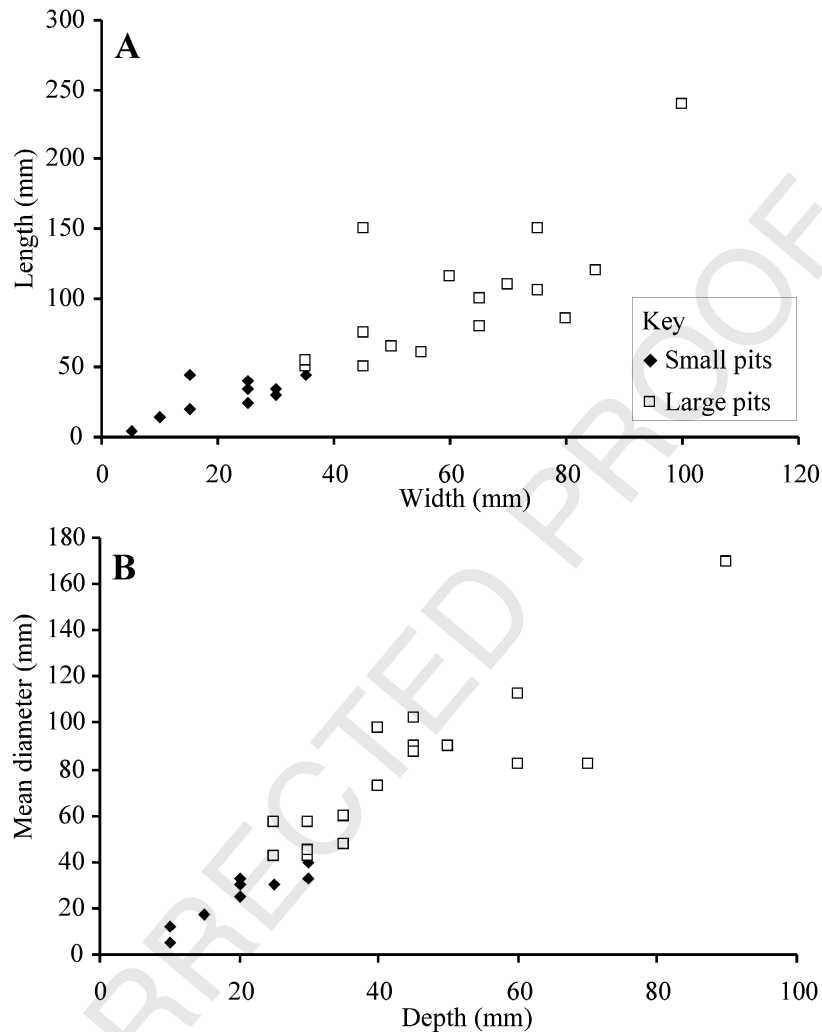


Fig. 4. Relationship between (A) elongation (length versus width) and (B) mean diameter versus depth of palaeokarstic depressions on surface 2.

256 compared to surface 2. The distinctive features of the
 257 surface are microkarren phenomena represented by
 258 small cylindrical potholes up to 40 mm in diameter and
 259 30 mm deep (Fig. 5A). These are similar to the rain-
 260 splash pits identified on surface 2. Two larger pits on
 261 surface 3 are about 90 mm in diameter and 45 mm deep.
 262 As with surface 2, dissolution extends into the body of
 263 the limestone bed producing a lensoid, open-joint
 264 appearance (Fig. 5B). The bedding in this rhythm is also
 265 steeply dipping and the palaeokarstic surface is normal
 266 to the steeply-inclined bedding.

267 Surface 3 is recognisable in all three sections of the
 268 Guinea Corn Formation studied (Fig. 2). In each case,
 269 the surface is overlain by a prominent limestone.

270 3. Interpretation of the surfaces

271 The microtopography of palaeokarstic surface 2 com-
 272 prises two principal karren types based on size, both of
 273 which would be described as circular plan forms in the

classification of Ford & Williams (1989), or as etched
 forms resulting from solutional attack in the types
 described by White (1988). Some of the larger features
 are solution pits, while others with detrital infill and flat
 floors with over-steepened and undercut bounding
 slopes are solution pans (solution basins); the latter were
 termed 'kamenitza' by Bögli (1960). Solution pits, to-
 gether with karren shafts, are the most widespread active
 karren forms on both bare limestone surfaces and be-
 neath a soil cover (Ford & Williams, 1989). In active
 karst environments, solution pans occur principally on
 bare or slightly vegetated rock, where they are a major
 karren form (Trudgill, 1985). These forms appear to be
 rare or absent beneath a soil cover, as solution pans
 develop where pools and ponds of water collect on the
 surface due to partial armouring of the pan floor by
 organic detritus or soil. This inhibits seepage into the
 limestone and focuses solution around the perimeter of
 the pan, over-steepening and undercutting its boundary
 slopes. In contemporary environments, solution pans

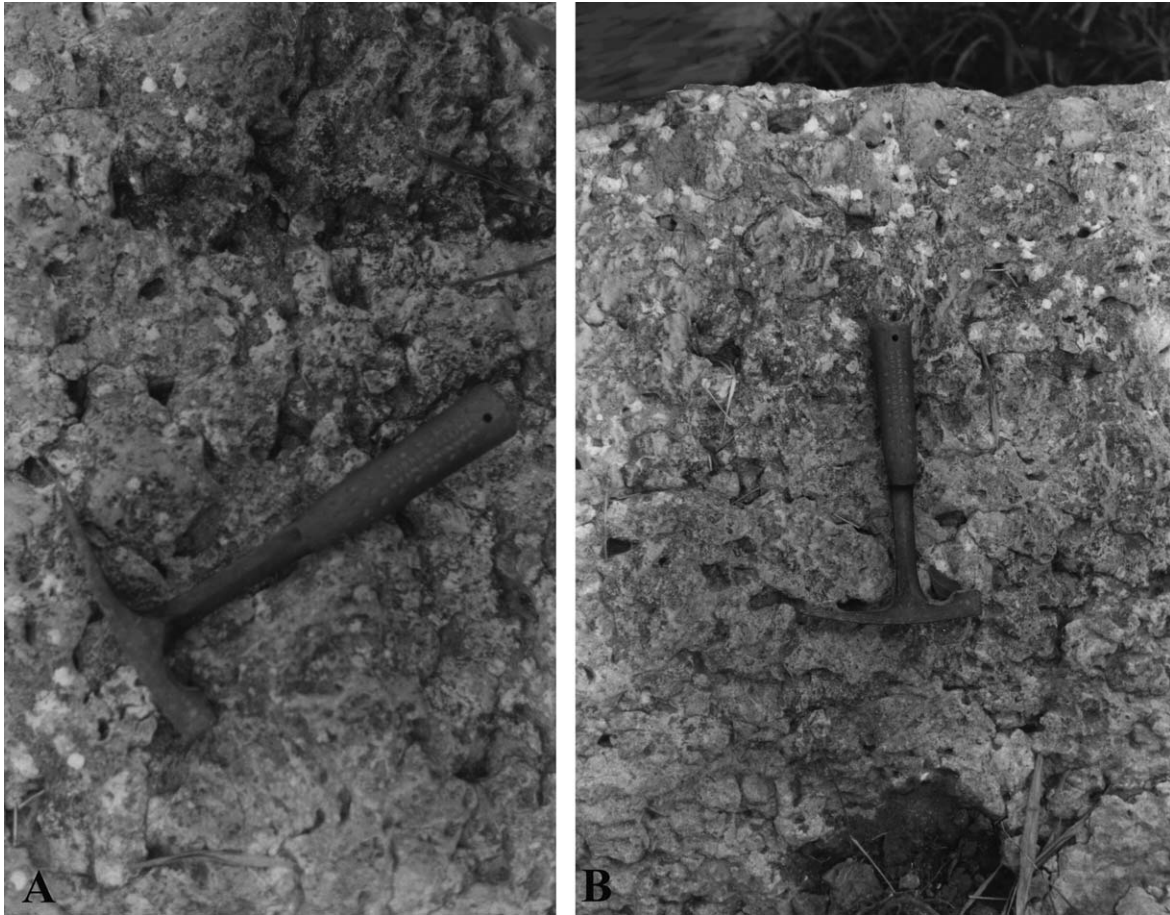


Fig. 5. A, mottled and pitted palaeokarstic surface 3 from cycle G5. The surface pitting is generally restricted to small pock marks. B, section of cycle immediately below the palaeokarstic surface. The limestone below the surface has been karstified to form an open lensoid fissure pattern.

294 and basins only occur on rock surfaces that are hori- 317
 295 zontal, such that water may collect in small pools and 318
 296 basins. 319

297 The smaller features recognised on surface 2 are also 320
 298 classified as pits in the terminology of Ford & Williams 321
 299 (1989), although some of the very small pits could 322
 300 alternatively be termed ‘rainpits’ *sensu* Jennings (1985). 323
 301 Rainpits are usually less than 30 mm in diameter and 324
 302 20 mm deep, often occurring in fields or singly when 325
 303 they are formed by leaf drip (Jennings, 1985). Such 326
 304 solution features in modern environments are also con- 327
 305 sidered to result from biogenic weathering, through the 328
 306 biochemical activity of cyanobacteria releasing CO₂ at 329
 307 night (Danin & Garty, 1983). 330

308 The larger potholes associated with surface 2 are 331
 309 analogous to modern-day karren shafts (Ford & 332
 310 Williams, 1989), which are short ‘caves’ up to 2–3 m 333
 311 deep and 1 m wide. Most have a structural guide and are 334
 312 formed by surface drainage into the epikarst. 335

313 Consequently, we ascribe the larger pits to dissolution 336
 314 associated with pools of water lying on the surface, 337
 315 since some of them are flat-floored solution pans, which, 338
 316 as indicated earlier, only occur on bare surfaces in 339

contemporary environments. The morphometry of the 317
 larger pits would suggest that, as solution progressed, 318
 they became larger, wider and deeper. Pools of standing 319
 water in the floor of the pits would also have led to 320
 solutional undercutting and over-steepening of their 321
 bounding slopes. We attribute the development of the 322
 small pock-marks to rain splash processes, suggesting a 323
 bare or lightly vegetated surface, either devoid of, or 324
 with an incomplete soil cover. The smaller pits are more 325
 isolated features, which may suggest that they formed 326
 from persistent at-a-point leaf drips and throughfall off 327
 vegetation onto a relatively soluble substrate. The two 328
 deeper potholes on surface 2 are attributed to more 329
 significant vadose water movements by strong surface 330
 drainage into the epikarst along more exploitable lines 331
 of weakness, suggesting at least partial induration and 332
 the development of a joint pattern. The solution scallop- 333
 ing and fluting on the interior walls of the potholes 334
 would also suggest the general lack of a soil cover on the 335
 exposure surface during the karstification phase. 336

The environmental setting of the surface at the 337
 time of karstification was one of a relatively soft sub- 338
 strate, possibly lightly vegetated, where small pits, with 339

ephemeral or permanent pools and puddles, developed by solution. Smaller pits developed on the surface by rain-splash processes in areas between the pools. The surface was sufficiently elevated to produce meteoric water movements from the surface through the vadose zone to the epikarst. The exposure of the carbonates to fresh-water led to meteoric diagenesis, when they became stabilised to low magnesium carbonates and sufficiently cemented to promote pitting of the surface. The dominance of solution over cementation in the formation of these surfaces suggests a humid tropical environment (Tucker & Wright, 1990).

Surface 2 at Cabbage Hill is overlain by a lignite rich siltstone. This deposit indicates high fluxes of land-derived terrestrial plant material. Such a scenario is consistent with transgressive flooding of a karstic terrain and the development of eutrophic conditions in shallow-water marine lagoonal areas (cf. Mitchell, 2002b).

4. Implications of exposure surfaces in the Guinea Corn Formation

Palaeokarstic surfaces have not previously been recorded from the Upper Cretaceous limestones of the Caribbean region. Johnson & Kauffman (1996) suggested that a brief period of emergence may have been responsible for the development of a hardground in the *Titanosarcolithes* limestones of the Jerusalem Mountain Inlier of western Jamaica; however, no palaeokarst was developed and submarine cementation seems more probable.

The existence of palaeokarstic surfaces in the Guinea Corn Formation capping offshore rudist-rich limestones implies uplift and not simply shoreline progradation. The tops of cycles generated through progradation of the shoreline do not display evidence of prolonged exposure, but are of fairly low relief and affected by saline or brackish water, which would preclude the generation of significant vadose alteration (Wright, 1994), in which case the cycle would not exhibit evidence of vadose meteoric overprinting in the form of karst solution features (Strasser, 1991). Small-scale palaeokarst of the nature described within the Guinea Corn Formation is best preserved where the karst processes are overwhelmed by rapid deposition, especially by terrigenous sediments or volcanic ash, succeeded by prolonged subsidence (Ford & Williams, 1989). Where palaeokarst is buried by a succeeding marine transgression, palaeokarstic surfaces may be expected to become destroyed by marine erosion or affected by bioerosion, particularly where wave energy is high. The palaeokarstic surfaces may, therefore, have been inundated by transgression in a low energy environment, though they are more likely to have become buried relatively

rapidly by estuarine organic and reworked volcanic sediments, forming the overlying beds.

The presence of two palaeokarstic surfaces suggests that there were at least two falls in sea level that exposed the limestones to karstification, and subsequently transgressions that allowed limestone deposition to resume. The lack of palaeokarstic surfaces capping other rhythms in the Guinea Corn Formation may be explained by transgressive erosion removing evidence of exposure, different water depths for different parts of the formation (i.e., relative falls in sea level were unable to expose the limestones), or by different degrees of relative sea level fall (sea-level falls were too small to expose the limestones). Small eustatic changes in sea level have been widely suggested as the cause of variations in stratal geometry in many Cretaceous successions (e.g., Haq et al., 1988; Plint, 1996). The Cretaceous is generally interpreted to have been a time of greenhouse climatic conditions with high atmospheric CO₂ levels (Crowley & Berner, 2001; Wilson et al., 2002). Yet despite this, the wide recognition of small-scale sea-level fluctuations has been interpreted to be due to the waxing and waning of small ice caps (e.g., Gale et al., 2002). Small-scale waxing and waning of these ice caps could have produced minor oscillations of sea level that would be necessary to produce the palaeokarst in the Guinea Corn Formation.

5. Conclusions

The origin of pock marks and pits on the top surfaces of sedimentary rhythms in the Guinea Corn Formation are interpreted as penecontemporaneous with the sedimentary succession. The karren features are considered to be palaeokarst for three reasons:

- 1 The pits and pots are perpendicular to bedding rather than to the present land surface. They are developed at the base of a rhythm that is steeply dipping, and there is a lack of any obvious asymmetry to their development. On a steeply-dipping bedding plane surface any later karren development along lines of weakness would be expected to produce asymmetric forms in both cross-section and plan, but this is not the case. Such a surface would also tend to develop hydrodynamically-controlled, linear karren features formed by sheet- or channel-flow, rather than the circular to oval etched forms that occur.
- 2 Some of the pock marks and small pots are infilled with very poorly sorted siltstone, which forms the base of the overlying rhythm. Also, the top 40 cm of the bed below the palaeokarstic surface in rhythm G1 is weathered into a honeycomb structure similar to spitzkarren, with an orientation normal to bedding and partly infilled with sedimentary rock from the overlying bed.

3 The surface was briefly exposed to vadose diagenesis and karren pitting followed by re-submergence and deposition of new carbonate above it. Thus, the surface fits into the definition of palaeokarst in that it was buried subsequent to a karstification phase which produced the small-scale solution features. It is also unrelated to any contemporary karst processes modifying the exposures today.

The palaeokarst in the Guinea Corn Formation suggests short periods of sea-level fall followed by sea-level rise. Such sea-level cycles have been suggested elsewhere and attributed to possible waxing and waning of small ice caps during the Late Cretaceous.

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